

# Logistic splicing correction for VNIR-SWIR reflectance imaging spectroscopy

FEDERICO GRILLINI<sup>1,\*</sup>, JEAN-BAPTISTE THOMAS<sup>1,2</sup>, AND SONY GEORGE<sup>1</sup>

<sup>1</sup> Coloumlab, Norwegian University of Science and Technology (NTNU), Norway

<sup>2</sup> ImViA Laboratory, University of Burgundy, France

\* Corresponding author: federico.grillini@ntnu.no

Compiled December 4, 2023

In the field of spectroscopy, a splicing correction is a process by which two spectra captured with different sensors in adjacent or overlapping electromagnetic spectrum ranges are smoothly connected. In our study we extend this concept to the case of reflectance imaging spectroscopy in the visible-near infrared (VNIR) and short-wave infrared (SWIR), accounting for additional sources of noise that arise at the pixel level. The proposed approach exploits the adaptive fitting of a logistic function to compute correcting coefficients that harmonize the two spectral sets. This short letter addresses usage conditions and compares results against the existing state of the art.

© 2023 Optica Publishing Group

<http://dx.doi.org/10.1364/ao.XX.XXXXXX>

In recent years many fields of research have experienced the deployment of reflectance imaging spectroscopy (RIS, also commonly known as hyperspectral imaging), often simultaneously combining the performances of imagers in the visible (VIS), near-infrared (NIR), and short-wave-infrared (SWIR) regions of the electromagnetic spectrum [1–5]. Albeit two different spectral sensors may capture the same physical quantity, namely spectral radiance, the reported values will hardly match if compared at corresponding wavelengths. The difference in response, which upon visual observation of two complementary spectra results in what has been defined in the literature as a "radiometric jump", "stepped data" or "spectral discontinuity", arises from a variety of factors that have been extensively studied by manufacturers in the field of spectroscopy [6, 7]. In [8], it was identified that the detector responsible for the observation of visible light was largely affected by warm-up time causing spectral sensitivity drift. In [7] the authors model the behavior of the VNIR sensor as a function of ambient temperature, while they find that the response of the infrared detector was not affected. One of the main reasons for the presence of discontinuities is the decreasing signal-to-noise ratio (SNR) at the extremities of the sensitive regions of the semiconductor materials from which the detectors are built: indeed, the absorption coefficients of the most com-

monly used semiconductors rarely overlap [9], making it hard to obtain a reliable combination of sensors in a wide range of wavelengths. For example, silicon (Si), the most used material to detect visible radiation, ends its operational range at around 1000 nm [10]. Materials such as Indium-Gallium-Arsenide (InGaAs) and Mercury-Cadmium-Telluride (HgCdTe) are used to detect infrared radiation, but need to be accurately designed (with an intrinsic concentration of elements) and appropriately cooled to be able to sense radiation at about 1000 nm [11]. Another factor that possibly generates jumps is the switch in bandwidth between two adjacent sensors, which leads to a different amount of energy incoming on the detectors, even for the same nominal wavelength.

Solutions to the problem of radiometric jumps are often referred to in the spectroscopy literature as "splicing correction" and include additive, multiplicative [12] and parabolic correction routines [13]. While the former two make use of a global correction scalar or coefficient, the latter solution proposes a wavelength-dependent coefficient to erase the radiometric jump by matching one spectrum to the other. In [7] the authors state that multiplicative and additive corrections lead to the introduction of more errors in the spectra, especially when considering high-energy spectra, while the parabolic method efficiently matches jumps up to a 6% difference, but struggles to correct larger spectral discrepancies.

In VNIR-SWIR RIS, although observed [14], the problem has not been deeply studied by the community. In remote sensing and airborne applications the bands around 970 nm are sometimes not processed due to the presence of a water absorption band [15], while with the motivation of low SNR, it is commonly accepted for laboratory applications to discard such flawed spectral bands and conduct further steps of analysis on disconnected or independent image datasets [1]. However, there is a need to preserve as much as possible of the available information to enhance visualization methods and to be able to highlight important spectral signatures in the interval of wavelengths that could be lost. When an imaging application is considered against a spectroscopy one, it is necessary to extend the list of factors that generate spectral discrepancies at the pixel level. First of all, the magnitude of the jumps is highly influenced by the performances of image registration, especially in the case in which the two images have different  $(x, y)$  dimensions and a scale difference exists. In this case, despite the fact that sub-pixel accuracy

can be achieved, the perfect pixel correspondence is an ideal condition. Moreover, reproducing the exact same relative positioning of object-illumination-camera is a rather difficult task when two separate imagers are deployed in an environment that can be adjusted only manually, and spurious differences in Bidirectional Reflectance Distribution Function (BRDF) [16] at the pixel level may generate a small difference in response.

The authors in [17] adapt the parabolic correction routine to concatenate VNIR and SWIR hyperspectral images of airborne sensing. In our previous attempt [18], global coefficients were learned through the optimization of a joint radiometric correction performed on standardized targets, but the results could not correctly generalize in terms of spectral variability and jump magnitude.

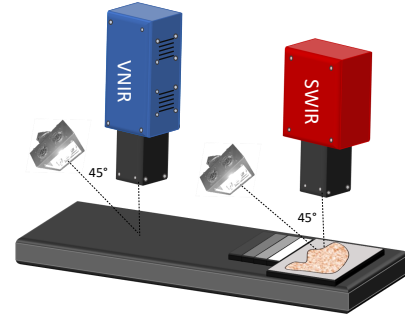
We propose a new adaptive splicing correction routine for complementary hyperspectral images that share an overlap of nominal wavelengths. Furthermore, we propose an evaluation that compares the results obtained with the proposed solution against the state of the art represented by the parabolic splicing correction. In this specific case, we consider a laboratory use of dual RIS in the VNIR and SWIR regions, deploying push-broom hyperspectral imagers manufactured by Hypspec (NEO, Norway), for which the main specifications are reported in Table 1. Upon inspection of the data, it is noticeable that the spectral ranges overlap in the region between 950 nm and 1000 nm with 16 and 9 bands for VNIR and SWIR respectively. The pre-processing of the images includes radiometric calibration and co-registration, adopting the methodology proposed in [19] and following the guidelines highlighted in one of our previous studies [20].

Name	VNIR	SWIR
Sensor	Si (CMOS)	HgCdTe
Cooling	NA	150 K
Spatial lines	1800	384
Spectral range [nm]	400 – 1000	1000 – 2500
FWHM [nm]	3.26	5.45

**Table 1.** Technical specifications of VNIR1800 and SWIR384 hyperspectral imagers.

The laboratory conditions of the set-up accentuate possible differences in BRDF due to the manual positioning of the illumination sources and the hyperspectral imagers. Thus, the experimental set-up (schematically reported in Figure 1) must be designed in a way that respects as much as possible the same illumination geometry for the two imagers. This also implies that the cameras are carefully aligned, both to maximize the overlap of the fields of view and to reduce pixel-wise differences of the BRDF. Furthermore, it will be necessary to calibrate the scene radiance captured by the two cameras and move into an illumination-independent space (absolute spectral reflectance), as there might be differences in the spectral power distribution (SPD) of the deployed illumination sources as well.

Let  $\mathbf{F}(x, y, \lambda)$  be a spectral image defined in the image domain  $\mathcal{D}$  with spatial coordinates  $(x, y)$  and spectral sampling in correspondence of wavelengths  $\lambda \in \Lambda$ . Generally,  $\Lambda$  is defined in  $[\lambda_{min}, \lambda_{max}]$ , but in our specific case it is the result of two separate image capture processes that generate



**Fig. 1.** Acquisition set-up of simultaneous push-broom VNIR-SWIR RIS. The illumination geometry (45/0) is carefully adjusted for both imagers. However, differences in SPD of the individual light source may exist.

$\mathbf{V}(x, y, \lambda_v)$  with  $\lambda_v \in \Lambda_v$ ,  $\Lambda_v = [400, 1000] \text{ nm}$  and  $\mathbf{S}(x, y, \lambda_s)$  with  $\lambda_s \in \Lambda_s$ ,  $\Lambda_s = [1000, 2500] \text{ nm}$ . Therefore, there exists a shared interval of wavelengths  $\Lambda' = \Lambda_v \cap \Lambda_s$ . Assuming that  $\mathbf{V}$  and  $\mathbf{S}$  are spatially co-registered and equally sampled in  $\Lambda'$ , we can define  $\mathbf{V}'(\lambda')$  and  $\mathbf{S}'(\lambda')$ . We aim at finding the transform  $\mathcal{T}$  that associates (dropping the spatial coordinates for readability):

$$\mathbf{F}(\lambda) = \mathcal{T}(\mathbf{V}(\lambda_v), \mathbf{S}(\lambda_s)) \quad (1)$$

For each image it is also possible to define the set of bands that lie outside of the overlapping range as  $\Lambda_v'' = \Lambda_v \setminus \Lambda'$  and  $\Lambda_s'' = \Lambda_s \setminus \Lambda'$ , associated to the *truncated* spectral images  $\mathbf{V}''(\lambda_v'')$  and  $\mathbf{S}''(\lambda_s'')$ .

The region in which  $\mathbf{V}'$  and  $\mathbf{S}'$  lie is where the radiometric jumps take place. Because of the aforementioned noise sources (different bandwidth, decreasing SNR, differences in BRDF, and sub-pixel misregistration), we assume that neither of the overlapping sets is a reliable estimate of the final spectrum. For this reason, we decide that a possible correct position could be the mean between  $\mathbf{V}'$  and  $\mathbf{S}'$ , noticing that this observation can be adjusted based on specific priors.

$$\mathbf{R} = \frac{1}{2} (\mathbf{V}' + \mathbf{S}') \quad (2)$$

Then we can define correcting coefficients  $\Phi'_v$  and  $\Phi'_s$  for VNIR and SWIR, valid in the overlapping region, such as

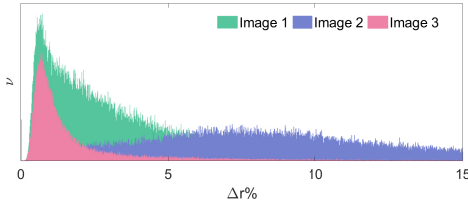
$$\mathbf{R} = \Phi'_v \mathbf{V}' = \Phi'_s \mathbf{S}' \quad (3)$$

When using the mean value as a reference for matching, the correcting coefficients will be found symmetrically distributed around 1. Ideally, the spliced spectrum should preserve the shapes of the original spectra and also match their magnitude values at some points away from the overlapping range. In order to do so, the correcting coefficients should smoothly vary from the values of  $\Phi'_v$  and  $\Phi'_s$  to 1. We achieve this by deploying a logistic function. In such a distribution, three parameters must be defined: the maximum value  $L$ , the slope  $k$ , and the center of the distribution  $x_0$ .

In this specific case,  $L$  is either the unity value or the coefficient at the extreme of the overlapping range:  $\varphi_v \in \Phi'_v$  (first coefficient) and  $\varphi_s \in \Phi'_s$  (last coefficient), while  $k$  and  $x_0$  are determined as a function of the distance  $\Delta r$  between  $\mathbf{V}'$  and  $\mathbf{S}'$ :

$$\Delta r = \sqrt{\frac{1}{N'} \sum_{i=1}^{N'} (\mathbf{V}' - \mathbf{S}')^2} \quad (4)$$

152 in which  $N'$  is the number of bands in the overlapping range. 186  
 153 In typical VNIR-SWIR applications, the values of  $\Delta r$  follow the 187  
 154 probability density function depicted in Figure 2. Here we can 188  
 155 observe that a value of 6% can already be considered very large.



156 **Fig. 2.** Typical relative spectral discrepancy histograms stem-  
 157 ming from three hyperspectral images.

156 The center of the logistic curve  $x_0$  is intrinsically linked to the  
 157 width of the spectral window that will experience the correction,  
 158 as it is the median value of the selected interval. Therefore,  $\lambda_0$   
 159 is obtained indirectly from the modeling of the window width  
 160  $w$ . To obtain smoothly connected spectra it is desirable to have a  
 161 steeper slope (higher  $k$ ) and a low  $w$  (number of spectral bands  
 162 affected by the correction) when the value of  $\Delta r$  is small, and vice  
 163 versa. The window width  $w$  is modeled as a logistic function of  
 164 the form:

$$165 \quad w_v = \frac{N_v''}{1 + \exp[-c_v(\Delta r - x_{0v})]} \quad (5)$$

$$166 \quad w_s = \frac{N_s''}{1 + \exp[-c_s(\Delta r - x_{0s})]} \quad (5)$$

165 in which  $N_v''$  and  $N_s''$  represent the number of bands in  $\Lambda_v''$   
 166 and  $\Lambda_s''$  respectively. The parameters  $c_v$ ,  $c_s$ ,  $x_{0v}$ , and  $x_{0s}$  are  
 167 learned by fitting the logistic function to the logarithmically  
 168 spaced values of  $\Delta r$ , in an interval that can be case-specific (in  
 169 the case of correction of pair of spectra) or empirically learned  
 170 from the  $\Delta r$  distribution (in the case images and large spectral  
 171 libraries). We can now define  $\Lambda_v''' \subset \Lambda_v''$  and  $\Lambda_s''' \subset \Lambda_s''$  as the  
 172 subsets that experience the splicing correction with a number of  
 173 bands equal to  $w_v$  and  $w_s$ , respectively. The central wavelengths  
 174  $\lambda_{0v}$  and  $\lambda_{0s}$  are then the median values of such intervals.

175 The slope  $k_v$  of the VNIR range can be found as the exponen-  
 176 tial function:

$$177 \quad k_v = a \cdot \exp(b\Delta r) + y \quad (6)$$

177 in which the parameters  $a$ ,  $b$ , and  $y$  are retrieved by fitting the  
 178 function to a linear vector decreasing from 1 to 0. We note here  
 179 that if the modeling is performed in the same way for the SWIR  
 180 counterpart, the normalization in  $[0,1]$  brings the fits of  $k_v$  and  
 181  $k_s$  to match. However, we can use prior information to model  
 182  $k_s$  so that it generates a flatter logistic curve since the affected  
 183 wavelength interval is larger.

$$184 \quad k_s = k_v \frac{N_v''}{N_s''} \quad (7)$$

184 The correction coefficients in  $\Lambda_v''$  and  $\Lambda_s''$  can now be deter-  
 185 mined as:

$$186 \quad \psi_v(\lambda_v'') = \frac{\text{sgn}(\varphi_v - \varphi_s)}{1 + \exp[-k_v(\lambda_v'' - \lambda_{0v})]} |\varphi_v - 1| + 1$$

$$187 \quad \psi_s(\lambda_s'') = \frac{\text{sgn}(\varphi_v - \varphi_s)}{1 + \exp[-k_s(\lambda_s'' - \lambda_{0s})]} |\varphi_s - 1| + \varphi_s \quad (8)$$

186 The full, smoothly connected hyperspectral image is obtained  
 187 by concatenating along the spectral dimension (dropping the  
 188  $\lambda$ -dependency for readability):

$$189 \quad \mathbf{F} = [\mathbf{V}'' \cdot \Psi_v, \mathbf{R}, \mathbf{S}'' \cdot \Psi_s] \quad (9)$$

189 An example of splicing correction on measured spectra can  
 190 be found in Figure S1 in Supplement 1.

191 The problem of spectral splicing has infinite solutions, as two  
 192 spectra can be connected and modified in infinite ways, but only  
 193 a limited set of pertinent solutions. Thus, the evaluation of the final  
 194 spectrum and the validation of the methodology can be tricky.  
 195 In order to help the evaluation, a Lambda1050 spectroradiometer  
 196 (Perkin Elmer inc.) was deployed to obtain a continuous ground  
 197 truth measurement in the interval 400 – 2500 nm. Such an in-  
 198 strument deploys two sensors: a Photomultiplier Tube (PMT)  
 199 for the range 400 – 860 nm and an Indium-Gallium-Arsenide  
 200 (InGaAs) detector for the range 861 – 2500 nm. Therefore, the  
 201 spectral region of interest of the VNIR-SWIR splicing correction  
 202 is included in a single sensor (InGaAs) sensitivity range. A total  
 203 of 175 samples coming from a collection of oil-painted mockups  
 204 [21] are measured. Such samples possess a level of texture high  
 205 enough to produce slight differences in BRDF at the pixel level.  
 206 In evaluating the final result, two properties of the reconstructed  
 207 spectra are evaluated:

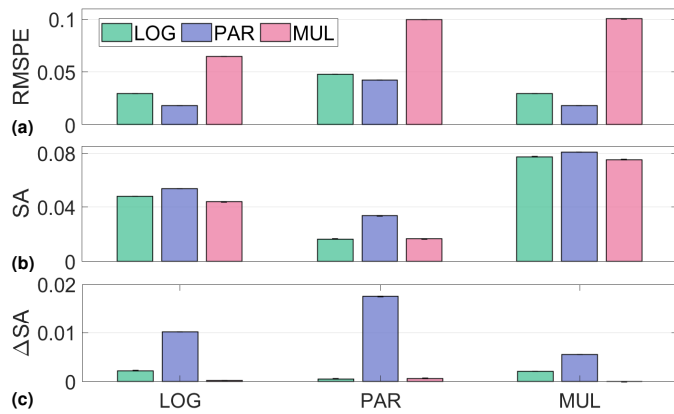
- 208 1. Conformity with the spectral shape of the ground truth,  
 209 evaluated through the usage of Spectral Angle (SA) [22],
- 210 2. Minimum intervention on the original spectra, measured by  
 211 means of the root mean square percentage error (RMSPE)  
 212 [23].

213 The rationale behind the choice of not considering a met-  
 214 ric that compares the absolute reflectance values of the ground  
 215 truth and the VNIR-SWIR spectra resides in the fact that the spec-  
 216 troradiometer averages the measurement over an area, while  
 217 the highly textural samples possess a high degree of spectral  
 218 variability that makes the magnitude comparison meaningless  
 219 at the pixel level. Furthermore, the acquisition geometries of  
 220 hyperspectral capture and spectroscopy are different.

221 The proposed method (LOG) is compared against the exist-  
 222 ing state of the art of splicing correction in spectroscopy. The  
 223 parabolic correction (PAR) was adapted following the insight  
 224 of [17] and correcting the last 60 bands of VNIR. The first VNIR  
 225 wavelength to be corrected is then 785 nm, while the juncture  
 226 point was selected in the middle of the overlap area at 973 nm.  
 227 Although discouraged from usage, we also include the multi-  
 228 plicative correction (MUL), computing the global coefficient as  
 229 the ratio between the SWIR and VNIR bands at 950 nm.

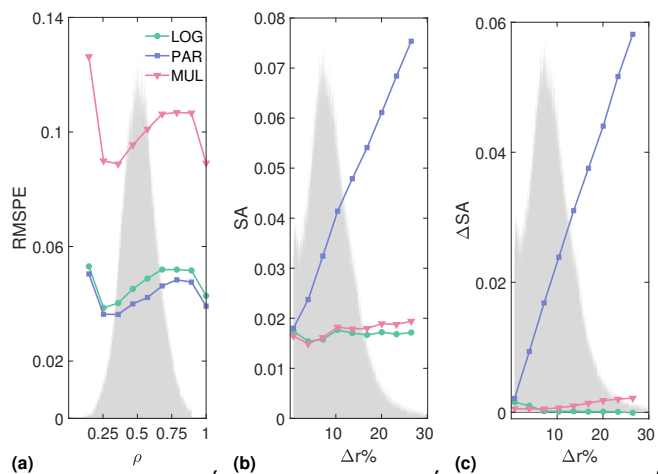
230 Since the selected correction methods affect a different num-  
 231 ber of bands, it is necessary to compare the evaluation metrics  
 232 in turn in the relative intervals of influence, as reported on the  
 233  $x$  axis of Figure 3. The results of SA between the spliced spec-  
 234 tra and the ground truth are however affected by the SA that  
 235 exists between the original disconnected spectra and the ground  
 236 truth. For this reason, it is decided to analyze the difference of  
 237 SA ( $\Delta SA$ ) in Figure 3c. Figure 3a highlights how the multiplica-  
 238 tive correction introduces a lot of unnecessary perturbations,  
 239 while the parabolic and logistic corrections achieve fairly sim-  
 240 ilar results. From Figure 3b and 3c it is possible to notice that  
 241 the proposed logistic correction produces more faithful spectral  
 242 shapes consistently.

243 Figure 4 highlights the limitations of the multiplicative and  
 244 parabolic corrections in some specific cases. When considering



**Fig. 3.** Mean values for  $RMSPE$  (a),  $SA$  (b), and  $\Delta SA$  (c) in different intervals of affected spectral bands: logistic (LOG), parabolic (PAR), and multiplicative (MUL).

a global coefficient for the whole spectrum, the relative magnitude perturbation that is introduced is highly impacting when the energy of the spectrum is low, as Figure 4a depicts. On the other hand, the parabolic correction shows a rapidly increasing perturbation in spectral shape as the spectral discrepancy also increases (Figures 4b and 4c), confirming the previous observation of [7]. The proposed logistic correction is proved to be more stable to such specific cases that have anyway a likely occurrence, as depicted by the frequency histograms of spectral energy at the overlap and relative spectral discrepancy.



**Fig. 4.** Behavior of the selected corrections in specific cases of spectral energy (a) and spectral discrepancy (b,c). The overlaid histograms (normalized for displaying purposes) illustrate the probability distributions of the events of spectral energy  $\rho$  and relative spectral discrepancy  $\Delta r\%$ .

In summary, we propose a new adaptive splicing correction routine to smoothly connect hyperspectral images that present spectral jumps in correspondence of adjacent spectral sensitivity intervals. The correction is performed in absolute reflectance space and it is adaptive in the sense that the amount of spectral bands affected depends on the magnitude of the initial spectral discrepancy. Advantages against the existing state of the art include better stability in cases of larger spectral discrepancies, which occur more likely in the case of imaging rather than in

point spectroscopy. The proposed method presents however a few shortcomings in splicing specific spectral shapes, as we highlight in Figure S2 and Figure S3 in Supplement 1. Therefore, it will be necessary in the future to extend the modeling of splicing correction to account for different levels of spectral complexity.

**Funding.** Placeholder for funding that will be automatically generated.

**Disclosures.** The authors declare no conflicts of interest.

**Data availability.** Data underlying the results presented in this paper are not publicly available at this time but may be obtained from the authors upon reasonable request.

**Supplemental document.** See Supplement 1 for supporting content. The code for splicing correction on individual spectra and on images can be found at <https://github.com/federigr/HyperspectralSplicingCorrection>.

## REFERENCES

- J. Sandak, A. Sandak, L. Legan, K. Retko, M. Kavčič, J. Kosel, F. Poohphajai, R. H. Diaz, V. Ponnuchamy, N. Sajinčič *et al.*, *Coatings* **11**, 244 (2021).
- A. Siedliska, P. Baranowski, M. Zubik, W. Mazurek, and B. Sosnowska, *Postharvest Biol. Technol.* **139**, 115 (2018).
- C. Camino, V. González-Dugo, P. Hernández, J. Sillero, and P. J. Zarco-Tejada, *Int. journal applied earth observation geoinformation* **70**, 105 (2018).
- M. Selva, B. Aiazzi, F. Butera, L. Chiarantini, and S. Baronti, *IEEE J. selected topics applied earth observations remote sensing* **8**, 3008 (2015).
- J. K. Delaney, J. G. Zeibel, M. Thoury, R. Littleton, M. Palmer, K. M. Morales, E. R. de La Rie, and A. Hoeningwald, *Appl. spectroscopy* **64**, 584 (2010).
- M. Danner, M. Locherer, T. Hank, and K. Richter, *GFZ Data Serv.* (2015).
- A. Hueni and A. Bialek, *IEEE J. Sel. Top. Appl. Earth Obs. Remote. Sens.* **10**, 1542 (2017).
- T. H. Hemmer and T. L. Westphal, *Algorithms for Multispectral, Hyperspectral, Ultraspectral Imag. VI* **4049**, 249 (2000).
- D. Benedikovic, L. Viro, G. Aubin, J.-M. Hartmann, F. Amar, X. Le Roux, C. Alonso-Ramos, É. Cassan, D. Marris-Morini, J.-M. Fédéli *et al.*, *Nanophotonics* **10**, 1059 (2021).
- B. Jalali and S. Fathpour, *J. lightwave technology* **24**, 4600 (2006).
- G. L. Hansen, J. Schmit, and T. Casselman, *J. Appl. Phys.* **53**, 7099 (1982).
- W. Dorigo, M. Bachmann, and W. Heldens, *User's manual, Ger. Aerosp. Cent. (DLR), Oberpfaffenhofen* (2006).
- D. Beal and M. Eamon, *Anal. Spectr. Devices, Inc* (2009).
- R. J. Murphy, S. T. Monteiro, and S. Schneider, *IEEE Trans. on Geosci. Remote. Sens.* **50**, 3066 (2012).
- J. A. Curcio and C. C. Petty, *JOSA* **41**, 302 (1951).
- F. O. Bartell, E. L. Dereniak, and W. L. Wolfe, *Radiat. scattering optical systems* **257**, 154 (1981).
- U. Okyay and S. D. Khan, *Photogramm. Eng. & Remote. Sens.* **84**, 781 (2018).
- F. Grillini, J.-B. Thomas, and S. George, *Color. Imaging Conf.* **2021**, 276 (2021).
- D. M. Conover, J. K. Delaney, and M. H. Loew, *Appl. Phys. A* **119**, 1567 (2015).
- F. Grillini, J.-B. Thomas, and S. George, 2022 12th Workshop on Hyperspectral Imaging Signal Process. Evol. Remote. Sens. (WHISPERS) pp. presented, awaiting proceedings (2022).
- F. Grillini, J.-B. Thomas, and S. George, *Sensors* **21**, 2471 (2021).
- F. A. Kruse, A. Lefkoff, J. Boardman, K. Heidebrecht, A. Shapiro, P. Barloon, and A. Goetz, *Remote. sensing environment* **44**, 145 (1993).
- H.-S. Lee and M.-T. Chou, *Int. J. Comput. Math.* **81**, 781 (2004).

## FULL REFERENCES

- 327
- 328 1. J. Sandak, A. Sandak, L. Legan, K. Retko, M. Kavčič, J. Kosel, F. Pooh-  
329 phajai, R. H. Diaz, V. Ponnuchamy, N. Sajinčič *et al.*, "Nondestructive  
330 evaluation of heritage object coatings with four hyperspectral imaging  
331 systems," *Coatings* **11**, 244 (2021).  
332 2. A. Siedliska, P. Baranowski, M. Zubik, W. Mazurek, and B. Sosnowska,  
333 "Detection of fungal infections in strawberry fruit by vnir/swir hyperspec-  
334 tral imaging," *Postharvest Biol. Technol.* **139**, 115–126 (2018).  
335 3. C. Camino, V. González-Dugo, P. Hernández, J. Sillero, and P. J. Zarco-  
336 Tejada, "Improved nitrogen retrievals with airborne-derived fluores-  
337 cence and plant traits quantified from vnir-swir hyperspectral imagery  
338 in the context of precision agriculture," *Int. journal applied earth obser-  
339 vation geoinformation* **70**, 105–117 (2018).  
340 4. M. Selva, B. Aiazzi, F. Butera, L. Chiarantini, and S. Baronti, "Hyper-  
341 sharpening: A first approach on sim-ga data," *IEEE J. selected topics  
342 applied earth observations remote sensing* **8**, 3008–3024 (2015).  
343 5. J. K. Delaney, J. G. Zeibel, M. Thoury, R. Littleton, M. Palmer, K. M.  
344 Morales, E. R. de La Rie, and A. Hoenigswald, "Visible and infrared  
345 imaging spectroscopy of picasso's harlequin musician: mapping and  
346 identification of artist materials in situ," *Appl. spectroscopy* **64**, 584–594  
347 (2010).  
348 6. M. Danner, M. Locherer, T. Hank, and K. Richter, "Spectral sampling  
349 with the asd fieldspec4," GFZ Data Serv. (2015).  
350 7. A. Hueni and A. Bialek, "Cause, effect, and correction of field spec-  
351 troradiometer interchannel radiometric steps," *IEEE J. Sel. Top. Appl.  
352 Earth Obs. Remote. Sens.* **10**, 1542–1551 (2017).  
353 8. T. H. Hemmer and T. L. Westphal, "Lessons learned in the postprocess-  
354 ing of field spectroradiometric data covering the 0.4-2.5-um wavelength  
355 region," in *Algorithms for Multispectral, Hyperspectral, and Ultraspectral  
356 Imagery VI*, vol. 4049 (SPIE, 2000), pp. 249–260.  
357 9. D. Benedikovic, L. Viro, G. Aubin, J.-M. Hartmann, F. Amar, X. Le Roux,  
358 C. Alonso-Ramos, É. Cassan, D. Marris-Morini, J.-M. Fédéli *et al.*,  
359 "Silicon–germanium receivers for short-wave-infrared optoelectronics  
360 and communications," *Nanophotonics* **10**, 1059–1079 (2021).  
361 10. B. Jalali and S. Fathpour, "Silicon photonics," *J. lightwave technology*  
362 **24**, 4600–4615 (2006).  
363 11. G. L. Hansen, J. Schmit, and T. Casselman, "Energy gap versus alloy  
364 composition and temperature in hg<sub>1-x</sub>cd<sub>x</sub>te," *J. Appl. Phys.* **53**,  
365 7099–7101 (1982).  
366 12. W. Dorigo, M. Bachmann, and W. Heldens, "As toolbox and processing  
367 of field spectra," User's manual, Ger. Aerosp. Cent. (DLR), Oberpfaf-  
368 fenhofen (2006).  
369 13. D. Beal and M. Eamon, "Preliminary results of testing and a proposal  
370 for radiometric error correction using dynamic, parabolic linear transfor-  
371 mations of "stepped" data," *Anal. Spectr. Devices, Inc* (2009).  
372 14. R. J. Murphy, S. T. Monteiro, and S. Schneider, "Evaluating classi-  
373 fication techniques for mapping vertical geology using field-based  
374 hyperspectral sensors," *IEEE Trans. on Geosci. Remote. Sens.* **50**,  
375 3066–3080 (2012).  
376 15. J. A. Curcio and C. C. Petty, "The near infrared absorption spectrum of  
377 liquid water," *JOSA* **41**, 302–304 (1951).  
378 16. F. O. Bartell, E. L. Dereniak, and W. L. Wolfe, "The theory and mea-  
379 surement of bidirectional reflectance distribution function (brdf) and  
380 bidirectional transmittance distribution function (btdf)," in *Radiation  
381 scattering in optical systems*, vol. 257 (SPIE, 1981), pp. 154–160.  
382 17. U. Okyay and S. D. Khan, "Spatial co-registration and spectral concate-  
383 nation of panoramic ground-based hyperspectral images," *Photogramm.  
384 Eng. & Remote. Sens.* **84**, 781–790 (2018).  
385 18. F. Grillini, J.-B. Thomas, and S. George, "Radiometric spectral fusion of  
386 vnir and swir hyperspectral cameras," in *Color and Imaging Conference*,  
387 vol. 2021 (Society for Imaging Science and Technology, 2021), pp.  
388 276–281.  
389 19. D. M. Conover, J. K. Delaney, and M. H. Loew, "Automatic registration  
390 and mosaicking of technical images of Old Master paintings," *Appl.  
391 Phys. A* **119**, 1567–1575 (2015).  
392 20. F. Grillini, J.-B. Thomas, and S. George, "Hyperspectral vnir-swir image  
393 registration: do not throw away those overlapping low snr bands," in  
394 *2022 12th Workshop on Hyperspectral Imaging and Signal Process-*  
*ing: Evolution in Remote Sensing (WHISPERS)*, (IEEE, 2022), pp.  
presented, awaiting proceedings.
- 395 21. F. Grillini, J.-B. Thomas, and S. George, "Comparison of imaging  
396 models for spectral unmixing in oil painting," *Sensors* **21**, 2471 (2021).  
397 22. F. A. Kruse, A. Lefkoff, J. Boardman, K. Heidebrecht, A. Shapiro,  
398 P. Barloon, and A. Goetz, "The spectral image processing system  
399 (sips)—interactive visualization and analysis of imaging spectrometer  
400 data," *Remote. sensing environment* **44**, 145–163 (1993).  
401 23. H.-S. Lee and M.-T. Chou, "Fuzzy forecasting based on fuzzy time  
402 series," *Int. J. Comput. Math.* **81**, 781–789 (2004).  
403  
404