

Development of frequency-dependent ocean wave directional distributions

Abushet Simanese^a, Harald Krogstad^b, Karsten Trulsen^{a,*}, José Carlos Nieto Borge^c

^a*Department of Mathematics, University of Oslo, PO Box 1053 Blindern, NO-0316 Oslo, Norway*

^b*Department of Mathematical Sciences, NTNU, NO-7491 Trondheim, Norway*

^c*Signal Theory and Communications Department, Superior Politechnic School, University of Alcalá, Alcalá de Henares, 28805, Madrid, Spain*

Abstract

The paper discusses the development of a frequency dependent directional spread from an initial condition of frequency-independence. The study applies basin directional measurements from the *Maritime Research Institute Netherlands* (MARIN), simulated data from a nonlinear wave equation and field measurements from the Ekofisk field. The basin experiments and numerical simulations are initialized with a JONSWAP spectrum with frequency-independent directional distributions. In both cases we observe the development of a strong frequency-dependence of the directional spread. The numerical simulations suggest that static nonlinear contributions to the surface elevation partially explain the behavior below the spectral peak in accordance with [1]. There are also dynamic nonlinear contributions on both sides of the spectral peak.

Keywords: Frequency dependent directional distribution

1. Introduction

In Gaussian linear wave theory (LWT), the ocean wave field is characterized by the directional spectrum

$$E(\omega, \theta) = S(\omega)D(\theta, \omega), \quad (1)$$

*Corresponding author

Email address: karstent@math.uio.no (Karsten Trulsen)

where S is the frequency spectrum, D the directional distribution, ω the angular frequency and θ is the direction [2]. If the required assumptions are made, the directional spectrum can easily be extracted from ocean wave data recorded using instruments ranging from the simplest single point triplet and gauge arrays to the more recent stereo video-based [3, 4, 5, 6] and remote-sensing [7] systems. Many studies dealing with the frequency dependent behavior of ocean wave directional distribution often assume the $\cos-2s$ distribution [8] as a primary choice, and express the spreading parameter, s , in terms of the dimensionless frequency to optimize the fit to measured directional characteristics. Proposed parameteric representations can be found from the recent studies of [9, 10, 11] and early work of [12, 13, 14] for practical applications. However, there are also several other idealized directional distributions with equivalent parameters as in the $\cos-2s$ distribution, suitable to describe the frequency dependence of the directional spread, or the circular standard deviation of D , [15]. A more comprehensive review of many of the parameterizations are given in [16]. Almost all analysis based on ocean wave data show that the directional spread is strongly dependent on frequency, with minimum spread near the spectral peak.

Since the pioneering work of Hasselmann [13], the frequency dependence of the angular distribution has been speculated to be due to the dominance of nonlinear interactions over generation effects. Second order spectral contributions derived from perturbation expansion of the surface elevation up to the 4th order in wave steepness [e.g. 17], reveal nonlinear influence on the shape of the directional spread. In particular, the difference-frequency contribution is the principal cause for the strong increase in the spread, and the discrepancy between the RMS and the linear wave theory wavenumber below the spectral peak [1]. The three-dimensional MNLS simulations of [18], carried out to study the establishment of high frequency power-law and the corresponding shift in the spectral peak, lead to the same conclusion.

We are not aware that basin wave spectra have yet been investigated whether they develop a frequency dependent directional spread on their own. However, in advanced basin wave makers, calibration of directional seas sometimes employs

realistic frequency dependent directional distributions [e.g. 19].

In the present paper, we take advantage of available experimental data to investigate the development of frequency dependent directional spread. The corresponding numerical simulations are done using the modified nonlinear Schrödinger equation [20, 21, 22, 23, 24].

We also present observations from the Ekofisk laser array, discuss data analysis methods, and observe bimodal nature of the directional distributions [25, 9, 26, 27].

The next section presents the theoretical background and analysis methods. Sec. 3 and 4 summarize data sources and results, respectively, followed by the conclusions in Sec. 5.

2. Background theory

2.1. Wave spectra

We consider a zero mean, weakly stationary and homogeneous ocean surface, $z = \eta(\mathbf{x}, t)$, where $\mathbf{x} = (x, y)$ is the horizontal position, t is time, and the z -axis pointing upwards. The spectral representation of η is a stochastic Fourier integral over the wavenumber-frequency (\mathbf{k}, ω) -space,

$$\eta(\mathbf{x}, t) = \int_{\mathbf{k}, \omega} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)} dZ(\mathbf{k}, \omega). \quad (2)$$

The Fourier transform in this context, Z , is called the spectral amplitude. From the properties of Z , the covariance function, $\rho(\mathbf{x}, t)$, and the spectrum of the surface are related by the integral

$$\rho(\mathbf{x}, t) = \mathbb{E}[\eta(\mathbf{x}, t)\eta(\mathbf{0}, 0)] = \int_{\mathbf{k}, \omega} e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)} d\chi(\mathbf{k}, \omega). \quad (3)$$

It is convenient to write $d\chi(\mathbf{k}, \omega) = S(\omega) \phi_\omega(\mathbf{k}) d\omega d\mathbf{k}$. Here $\phi_\omega(\mathbf{k})$ is the wavenumber distribution at frequency ω , normalized so that $\int_{\mathbf{k}} \phi_\omega(\mathbf{k}) d\mathbf{k} = 1$ for all $\omega \in \mathbb{R}$, where $d\mathbf{k} \equiv dk_x dk_y$ is a short-hand notation. In the simplest case with only linear free waves, \mathbf{k} and ω are connected by the dispersion relation, e.g. $\omega = \sqrt{g|\mathbf{k}| \tanh(|\mathbf{k}|h)}$, $k_x = |\mathbf{k}| \cos \theta$, $k_y = |\mathbf{k}| \sin \theta$, and ϕ_ω may be

written as $\phi_\omega(\mathbf{k}) = \delta(|\mathbf{k}| - k_{LWT}(\omega)) D(\theta, \omega)$, where k_{LWT} is the solution of the dispersion relation for a fixed ω , and $D(\theta, \omega)$ is the angular distribution located at the intersection of the dispersion manifold and the plane $\omega = \text{const}$.

Following the discussion in [2], $\phi_\omega(\mathbf{k})$ will in general have contributions also off the dispersion manifold. Being a two-dimensional distribution, it is reasonable to consider its basic moments up to second order,

$$\langle (k_x)^i (k_y)^j \rangle(\omega) = \int_{\mathbf{k}} (k_x)^i (k_y)^j \phi_\omega(\mathbf{k}) d\mathbf{k}, \quad 0 \leq i + j \leq 2. \quad (4)$$

Apart from rather crude large-area remote-sensing systems, direct field measurements of $\phi_\omega(\mathbf{k})$ are out of reach. However, estimates of these moments are possible from single point surface heave/pitch/roll (HPR) systems. The method is well established, but unfortunately, the leading five moments are far from determining the actual shape of ϕ_ω , as illustrated by the following two possible solutions. The first solution is a two-dimensional Gaussian distribution, whereas the second is a quasi-LWT solution where the wave number is estimated from the data without enforcing by the dispersion relation [2, 28]. The first has hardly been used, but the second has become a standard for heave/pitch/roll systems where four Fourier coefficients of $D(\theta, \omega)$ may be uniquely obtained from the moments.

The preferred object in research, in particular in remote sensing is the wavenumber spectrum

$$\Psi(\mathbf{k}) d\mathbf{k} = 2 \int_{d\mathbf{k}, \omega > 0} d\chi(\mathbf{k}, \omega) = 2 \int_{\omega > 0} S(\omega) \phi_\omega(\mathbf{k}) d\mathbf{k} d\omega. \quad (5)$$

However, in the general case, $\Psi(\mathbf{k})$ will contain contributions from several ϕ_ω -distributions for each \mathbf{k} , and there is no way to discriminate between contributions from linear and nonlinear waves without further information about the ϕ_ω -functions. Inversion from Ψ back to χ is thus *only* possible assuming the dispersion relation, such as $\sigma(|\mathbf{k}|) = \sqrt{g|\mathbf{k}| \tanh(|\mathbf{k}|h)}$,

$$d\chi(\mathbf{k}, \omega) = \frac{1}{2} [\Psi(\mathbf{k}) \delta(\omega - \sigma(|\mathbf{k}|)) + \Psi(-\mathbf{k}) \delta(\omega + \sigma(|\mathbf{k}|))] d\mathbf{k} d\omega. \quad (6)$$

The transformation from the wavenumber spectrum to the commonly used

directional spectrum, $E(\omega, \theta) = S(\omega) D(\theta, \omega)$, is carried out by change in variables:

$$E(\omega, \theta) = S(\omega) D(\theta, \omega) = \Psi(|\mathbf{k}|(\omega), \theta) |\mathbf{k}|(\omega) \frac{d|\mathbf{k}|(\omega)}{d\omega}. \quad (7)$$

It is convenient to normalize S and Ψ so that

$$\text{Var}[\eta] = \int_{\mathbf{k}, \omega} d\chi(\mathbf{k}, \omega) = \int_{\mathbf{k}} \Psi(\mathbf{k}) d\mathbf{k} = 2 \int_0^\infty S(\omega) d\omega. \quad (8)$$

The angular distribution function $D(\theta, \omega)$ may be viewed as a probability distribution function distributed over the direction θ of \mathbf{k} , often stated in terms of its Fourier series expansion,

$$D(\theta, \omega) = \frac{1}{2\pi} \left[1 + 2 \sum_{n=1}^{\infty} [a_n(\omega) \cos n\theta + b_n(\omega) \sin n\theta] \right]. \quad (9)$$

The main parameters of the distribution are expressed in terms of the Fourier coefficients $a_1(\omega)$ and $b_1(\omega)$ [see 29, 2]. The mean direction is given by $\theta_1(\omega) = \text{atan2}(b_1(\omega), a_1(\omega))$, whereas the directional spread, $\sigma_1(\omega) = \sqrt{2(1 - r_1(\omega))}$, where $r_1(\omega) = \sqrt{a_1(\omega)^2 + b_1(\omega)^2}$. The $\cos-2s$ distribution is a favored choice in modeling directional wave fields:

$$D(\theta) = D_s(s) \cos^{2s} \frac{(\theta - \theta_1)}{2}, \quad D_s(s) = \frac{1}{\pi} 2^{2s-1} \frac{\Gamma^2(s+1)}{\Gamma(2s+1)}, \quad \sigma_1 = \sqrt{\frac{2}{s+1}}, \quad (10)$$

here s can be constant or in accordance with one of the empirical frequency dependent forms proposed by e.g. [14, 13] or [9]. For example, the Mitsuyasu parameterization [14] takes the form

$$s(\omega) = \begin{cases} s_p \left(\frac{\omega}{\omega_p}\right)^5 & \text{if } \omega \leq \omega_p, \\ s_p \left(\frac{\omega}{\omega_p}\right)^{-2.5} & \text{if } \omega > \omega_p, \end{cases} \quad (11)$$

65 where $s_p = 11.5(U/c_p)^{-2.5}$, U is the wind speed and c_p the phase speed at the spectral peak.

It is possible to relate the Fourier coefficients directly to the moments of ϕ_ω -functions, [2]. In particular, introducing the mean wavenumber, $\langle \mathbf{k} \rangle(\omega) =$

$\langle(k_x, k_y)\rangle$, and RMS wavenumber, $k_{RMS}(\omega) = \langle k_x^2 + k_y^2 \rangle^{1/2}$, we obtain the general expressions

$$\theta_1(\omega) = \angle \mathbf{r}(\omega), \quad \sigma_1(\omega) = \sqrt{2(1 - |\mathbf{r}(\omega)|)}, \quad \mathbf{r}(\omega) = \frac{\langle \mathbf{k} \rangle(\omega)}{k_{RMS}(\omega)}. \quad (12)$$

These definitions can be used independently of any additional assumptions beyond stationarity and homogeneity. We also employ the useful Check Ratio, $CR(\omega) = \frac{k_{RMS}(\omega)}{k_{LWT}(\omega)}$, where $k_{LWT}(\omega)$ is obtained from the dispersion relation.

70 2.2. Methods of data analysis

Stochastic wave data analysis mostly applies Fourier-based methods. This is ideal for Gaussian linear wave theory, where any linear and shift invariant filter, $\eta(\mathbf{x}, t) \rightarrow \xi(\mathbf{x}, t)$, may be expressed as

$$\xi_m = \int_{\mathbf{k}, \omega} T_m(\mathbf{k}, \omega) e^{i(\mathbf{k} \cdot \mathbf{x}_m - \omega t)} dZ(\mathbf{k}, \omega). \quad (13)$$

Here T_m is the transfer function and $e^{i\mathbf{k} \cdot \mathbf{x}_m}$ the location exponential. A first step in the analysis, following the data inspection and restoration, is to estimate the auto- and cross-spectral matrix, $\mathbf{\Sigma}(\omega) = [\sigma_{mn}]_{m,n=1}^M$, defined by

$$\sigma_{mn}(\omega) = \int_{\mathbf{k}} T_m(\mathbf{k}, \omega) T_n^H(\mathbf{k}, \omega) e^{i\mathbf{k} \cdot (\mathbf{x}_m - \mathbf{x}_n)} d\chi(\mathbf{k}, \omega), \quad m, n = 1, \dots, M, \quad (14)$$

where H stands for Hermitian transposed [2]. The transfer functions in the integral Eq. 14, including the location exponentials may be expressed in terms of ω and θ as $T_i(\omega, \theta) = R_i(\omega) h_i(\omega, \theta)$, where $R_i(\omega)$ and $h_i(\omega, \theta)$ contain the direction-independent and dependent parts, respectively [30]. If we further assume a directional spectrum of the form given in Eq. 7, then Eq. 14, may be rewritten as

$$\mathbf{\Sigma}(\omega) = S(\omega) \mathbf{R}(\omega) \mathbf{\Phi}(\omega) \mathbf{R}^H(\omega), \quad \mathbf{R} = \text{diag}(R_1, \dots, R_N), \quad (15)$$

where

$$\mathbf{\Phi}(\omega) = \int_0^{2\pi} \mathbf{h}(\omega, \theta) \mathbf{h}^H(\omega, \theta) D(\theta, \omega) d\theta. \quad (16)$$

We refer to [2] for computational details, but recall that estimation of $\mathbf{\Sigma}$ is made by averaging the products of the discrete Fourier transform of the time

series provided the data are well-sampled. This means that the data records are considerably longer than the temporal correlation, and the sampling frequency is well above the most energetic parts of the spectrum. In addition, data windowing and spectral averaging are used for decreasing spectral leakage and enhance the quality of the estimate.

There are many different estimation techniques for wave spectra, each having their own principle of analysis and merits [31]. We consider the estimation of directional wave spectra from spatial arrays of wave gauges applying the maximum likelihood (ML) method, the iterative maximum likelihood (IML) method, the heave/pitch/roll (HPR) method, and directly from $\chi(\mathbf{k}, \omega)$ when feasible. The ML estimate has the form [32]

$$D_{ML}(\theta, \omega) = \frac{\kappa}{\mathbf{h}^H(\omega, \theta)\mathbf{\Phi}^{-1}(\omega)\mathbf{h}(\omega, \theta)}, \quad (17)$$

where

$$\kappa^{-1} = \int_0^{2\pi} (\mathbf{h}^H(\theta, \omega)\mathbf{\Phi}^{-1}\mathbf{h}(\theta, \omega))^{-1} d\theta. \quad (18)$$

Data redundancy may sometimes require use of generalized inverses. However, cross-spectra computed from a certain $D(\theta)$ and then used to compute D_{ML} from Eq. 17 do not always match. Therefore, Pawka [33] introduced an iterative improvement of ML method called iterative maximum likelihood (IML) method, later modified to an iterative scheme of the form

$$D_{n+1} = D_n + \omega_R[\hat{D}_{ML} - M(D_n)] \quad \text{with} \quad D_0 = \hat{D}_{ML}. \quad (19)$$

In Eq. 19, \hat{D}_{ML} is the ML estimate based on the data, $M(D_n)$ the ML estimate based on cross-spectrum of \hat{D}_{ML} , ω_R a relaxation parameter, and D_{n+1} is normalized to have integral equal 1, [32, 34].

One may alternatively analyze three sensors as HPR data by adding a plane through the three measurements and approximate the surface elevation, η , and slopes, $\frac{\partial\eta}{\partial x}$ and $\frac{\partial\eta}{\partial y}$ at the center of gravity. In this case, we get single-point triplet measurements which may be analyzed using the standard buoy processing techniques [see, 35, p.184], leading to, among others, the Burg [36] maximum

entropy (ME-B) and the Shannon [37] maximum entropy (ME-S) directional distributions.

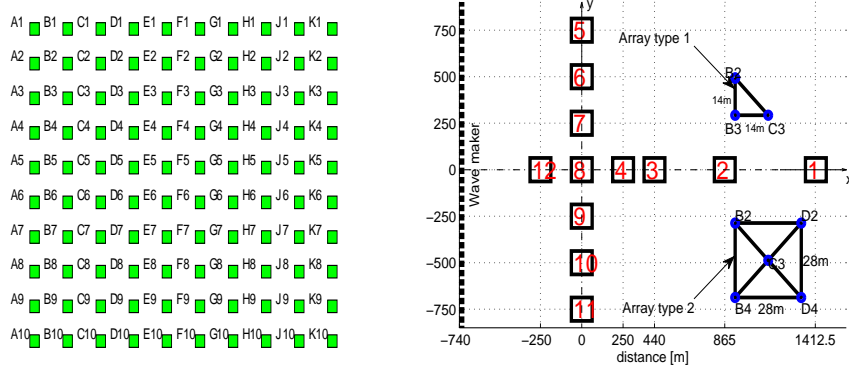
3. The different data sources

3.1. Laboratory measurements

90 Data for this study were collected by MARIN as part of the On board Wave and Motion Estimator (OWME) project. The experiment was conducted between Dec. 10 to Dec. 13, 2007, and was initiated with the main objective of developing an on board decision support system for vessel motions prediction. A detailed description of the facilities and experiments is presented in [38].

95 Model experiments at a scale of 1 : 70 were performed in the 40m long and 170m wide MARIN seakeeping and manoeuvring basin. The wave measurements were carried out using a 1.8m by 1.8m 10×10 wave probe array of grid size 0.2m in both directions as shown in Fig. 1a. The array was connected to a movable carriage which could be moved to various locations in the basin between
100 tests, each such location will in the following be referred to as array location. Arrangement of the wave gauge arrays, as well as their layouts throughout the 5m deep (350m in full scale) model basin is shown in Fig. 1b, the main wave propagation direction is along the positive x -axis. The waves are generated by a flap type wave maker at one end and damped by an absorbing beaches on the
105 opposing end. The wave maker is located -740m from the y -axis.

The wave maker was programmed with a JONSWAP spectrum with a directional distribution independent of frequency. The JONSWAP spectrum was generated with peakedness factor $\gamma = 3.3$, significant wave height $H_s = 2.5$ m, a peak period of $T_p = 9$ s. Three different directional distributions were employed,
110 narrow ($\sigma_{1I} = 5.7^\circ$), medium ($\sigma_{1I} = 12.5^\circ$) and broad ($\sigma_{1I} = 19.1^\circ$). The directional distribution was similar to the standard cos-2s-distribution introduced, see [8]. The spectral peak frequency in the measurement was $f_p = 0.113$ Hz, corresponding to $k_p = 0.0514$ rad/m, wavelength $\lambda_p = 122$ m and dimensionless



(a) Resistance type wave probes arranged in a square with sides 126 metres (full scale). (b) Full scale wave gauge array locations. The distance between the wave maker and the center of location 12 is 490m.

Figure 1: Wave gauge array locations and probe arrangement

depth $k_p h = 19$. In the following, we define the steepness ϵ as $\epsilon = k_p \sqrt{2\langle \eta^2 \rangle}$,
 115 accordingly the steepness in MARIN data is found to be $\epsilon = 0.05$.

Analysis results presented in Sec. 4.1 are based on measurements either from
 array type 1 or from array type 2. The former formed by probes B2, B3 and
 C3, while the latter formed by probes B2, B4, C3, D2 and D4 as depicted in
 Fig. 1b. The target location in the basin will be indicated in the caption when
 120 needed. The noise level in the auto-spectra is about 40dB below the peak and
 limits the useful part of the spectra to an interval, $[0, 4\omega_p]$.

Due to some noise of unknown origin, neighboring data series show lack of
 coherence in the cross-spectra above twice the spectral peak. The noise in the
 individual time series implies a fast drop in coherence in the cross-spectrum and
 125 a corresponding poor quality of the estimated wave parameters. In addition,
 data from manageable subarrays from the full arrays suffer from spatial aliasing
 and limits the usable frequency range for the directional analysis even further.
 In the MARIN probe set-up, the smallest leg in the array is 14m, resulting in
 a limiting wavelength of 28m. The corresponding frequency is 0.24Hz, or about

130 two times the spectral peak frequency.

3.2. Field measurements

The Ekofisk laser array consists of four down-looking OptechTM lasers mounted on a bridge connecting the Kilo and Bravo platforms near the Ekofisk complex. The lasers are placed at the four corners of a 2.6m by 2.6m square
135 located approximately 20 meters above the mean surface of a 70m deep sea.

The system has been designed by the Norwegian Meteorological Institute in cooperation with the University of Miami, and the data collection is carried out under the operational responsibility of ConocoPhillips Inc. The lasers use continuous 5Hz sampling frequency which was down-sampled to 1.7Hz for time
140 series of duration 20 minutes [39]. Sophisticated data check and restoration techniques have been implemented on the raw data [40]. The present analysis is limited to records with moderate to high sea states, that is, $H_s \geq 3.5\text{m}$. The spectral peak frequency is found to be $f_p = 0.12\text{Hz}$, the dispersion relation gives $k_p = 0.0616\text{rad/m}$, $\lambda_p = 2\pi/k_p = 102\text{m}$ and thus $k_ph = 4.3$.

145 Even if the full array consists of four lasers, the results are virtually unchanged when using only three lasers because of the array's compact size. The shortest leg in the array is 2.6m, resulting in a limiting wavelength $\lambda_{SA} = 5.2\text{m}$. The corresponding frequency, $f_{SA} \approx 0.56\text{Hz} = 4.4f_p$, which is well below the temporal Nyquist limit $f_s = 0.85\text{Hz}$.

150 3.3. Numerical model

We shall employ a modified nonlinear Schrödinger model that describes the evolution of the surface elevation using the evolution of its complex envelope, B . In this section all expressions are non-dimensionalized by a characteristic wavenumber (k_c) and the corresponding characteristic angular frequency (ω_c). In practice we let the characteristic frequency be the peak frequency of the initial spectrum. Assuming a two dimensional, irrotational flow of an incompressible inviscid fluid, we expand the velocity potential ϕ and surface elevation η of the

free surface in harmonic series

$$\phi = \bar{\phi} + \frac{1}{2}(Ae^{i(x-t)} + A_2e^{2i(x-t)} + A_3e^{3i(x-t)} + \dots + c.c.), \quad (20)$$

and

$$\eta = \bar{\eta} + \frac{1}{2}(Be^{i(x-t)} + B_2e^{2i(x-t)} + B_3e^{3i(x-t)} + \dots + c.c.). \quad (21)$$

Here *c.c.* denotes the complex conjugate, (x, y) and z are horizontal and vertical coordinates, t is time and $\bar{\eta}$, B_2 and B_3 are the zeroth, second and third harmonic bound waves. They are slowly varying functions of space and time given by

$$\bar{\eta} = -\bar{\phi}_t, \quad B_2 = \frac{1}{2}B^2 + iBB_t, \quad \text{and} \quad B_3 = \frac{3}{8}B^3. \quad (22)$$

Detailed representations of temporal and spatial Schrödinger equations are given in [23]. The temporal evolution equations have nice properties to investigate temporal sequences of measurements obtained from large-area remote sensing systems. However, for a meaningful study of the spatial development
 155 of frequency dependent directional spread away from a wave maker in a basin it is desirable to employ spatial MNLS equation. In particular, when the measurements are in the form of simultaneously recorded time series of certain wave properties taken at different spatial locations, the spatial MNLS equation is more suitable than the temporal MNLS. Moreover, conventional analysis meth-
 160 ods often use time series to estimate the basic directional parameters.

Provided that the steepness is small $\epsilon = k_c a_c \ll 1$, the bandwidth is narrow $|\Delta \mathbf{k}|/k_c \ll 1$, large depth $(k_c h)^{-1} \ll 1$, and neglecting terms of relative order of magnitude higher than ϵ^4 , we get the modified nonlinear Schrödinger (MNLS) equation. Here $a_c = \sqrt{2\langle \eta \rangle^2}$ is the characteristic amplitude, $|\Delta \mathbf{k}|$ is a characteristic modulation wave vector relative to the characteristic wave vector, [21] and h is depth set to infinity for the present numerical simulations. The potential $\bar{\phi}$ of the induced mean flow is governed by the equations

$$\begin{aligned} \bar{\phi}_z &= -(|B|^2)_t \quad \text{at} \quad z = 0, \\ 4\bar{\phi}_{tt} + \bar{\phi}_{yy} + \bar{\phi}_{zz} &= 0 \quad \text{for} \quad -\infty < z < 0, \\ \bar{\phi}_z &= 0 \quad \text{at} \quad z \rightarrow -\infty. \end{aligned} \quad (23)$$

The first, second and third harmonic complex amplitudes of the potential, A , A_2 and A_3 , will not be considered here. The slowly varying functions $\bar{\eta}$ and $\bar{\phi}$ contain the difference-frequency contributions. The modified nonlinear Schrödinger equation [41, 42] is the lowest order that can capture the behavior observed in [1]. The spatial evolution form of MNLS equation is given by [20, 21, 24]

$$B_x + 2B_t + iB_{tt} - \frac{i}{2}B_{yy} + i|B|^2B - B_{t yy} - 8|B|^2B_t - 2B^2B_t^* - 4i\bar{\phi}_t B = 0, \quad (24)$$

superscript $*$ is complex conjugate. We employ the numerical method of [43, 44] with periodic boundary conditions in time and transverse direction. Evolution in the x -direction is achieved with a splitting scheme in which the linear part of Eq. 24 is integrated exactly in Fourier space and the nonlinear part is integrated
165 by finite differences. We employ temporal and transversal grids with $N_t = 512$ and $N_y = 256$ points for simulation of B . The computational domain is set to 200 peak periods in time and 100 peak wavelengths in transversal direction with discretization $\Delta\omega = 1/200$ and $\Delta k_y = 1/100$. However, a different grid size namely, $N_t = 4096$ and $N_y = 2048$ points, is employed for the reconstruction of
170 the surface elevation using Eq. 21, in this case we set the discretization of the (t, y) -plane to be $\Delta t = \Delta y = 0.3$. For the directional analysis, evolution results are extracted at ten positions for every $10\lambda_c$ interval over a total length of about $100\lambda_c$ in the x -direction. The computational efforts range from an hour to few hours on office computers.

All simulations are started by a JONSWAP frequency spectrum with $\gamma = 3.3$, multiplied by directional distributions of the form [2]

$$D(\theta) = D_k(k) \cos^k(\theta - \theta_1), \quad \text{for } |\theta - \theta_1| \leq \frac{\pi}{2}, \quad (25)$$

where $D_k(k)$ is a normalization factor. This distribution is slightly advantageous for MNLS-type simulations in comparison with $\cos-2s$, because it does not prescribe waves going backward. We recall that $\sigma_1 = \sqrt{2/(s+1)}$ for the $\cos-2s$ -distribution, on the contrary there is no simple analytic expression for k as a function of σ_1 . Nevertheless, since the distribution in Eq. 25 and the

cos-2s-distribution are practically indistinguishable when $k = (s - 1)/2$ is larger than about 5 [2], one may use the approximation

$$k = (s - 1)/2 = \frac{1}{\sigma_1^2} - 1, \quad (26)$$

175 for $\sigma_1 < 1$.

The spectra shown below are obtained by analysing a 3D surface data $\eta(x, y, t)$, from MNLS simulation. The squared modulus of the *3D Fast Fourier Transform (3D-FFT)* of the data is a 3D wave spectrum $\chi(\mathbf{k}, \omega) = |\hat{\eta}(\mathbf{k}, \omega)|^2$, often interpreted as the distribution of wave energy in the (\mathbf{k}, ω) -space. Spectral leakage has been reduced by applying a cosine-bell tapering window on the
180 data. Besides, all the spectra have been smoothed with a 3D moving average.

The $k_x\omega$ -slice of $\chi(\mathbf{k}, \omega)$ at $k_y = 0$, is shown in Fig. 2. It displays the spectrum of the free waves including the zeroth harmonic and higher order nonlinear harmonics.

185 The wavenumber distribution, $\phi_\omega(\mathbf{k})$, has extracted from $\chi(\mathbf{k}, \omega)$ and shown in Fig. 3, for a set of positive frequencies. It demonstrates the $k_x k_y$ -slices carrying the \mathbf{k} -distributions near the most energetic parts of the zeroth, first, second and third order spectra. The zeroth harmonic spectrum has localized itself near the origin while the first order spectrum dominates the region around the spectral peak. It fulfills the linear dispersion, $|\mathbf{k}| = k_{lwt}(\omega)$ (white solid circle).
190 The 2nd and 3rd order spectra dominate the region around twice and three times the peak frequency respectively. They fulfill $|\mathbf{k}| = 1/2 k_{lwt}(\omega)$ (yellow dash-dot circle) and $|\mathbf{k}| = 1/3 k_{lwt}(\omega)$ (red dashed circle) respectively.

4. Results

195 4.1. Laboratory observations

Full 3-D spectra were obtained from three-dimensional discrete Fourier transform in space and time of *MARIN*'s 10×10 -array data. To reduce spectral leakage, the Fourier transform has been tapered with a cosine-bell window. In addition, we have applied a three-dimensional moving average to smooth the

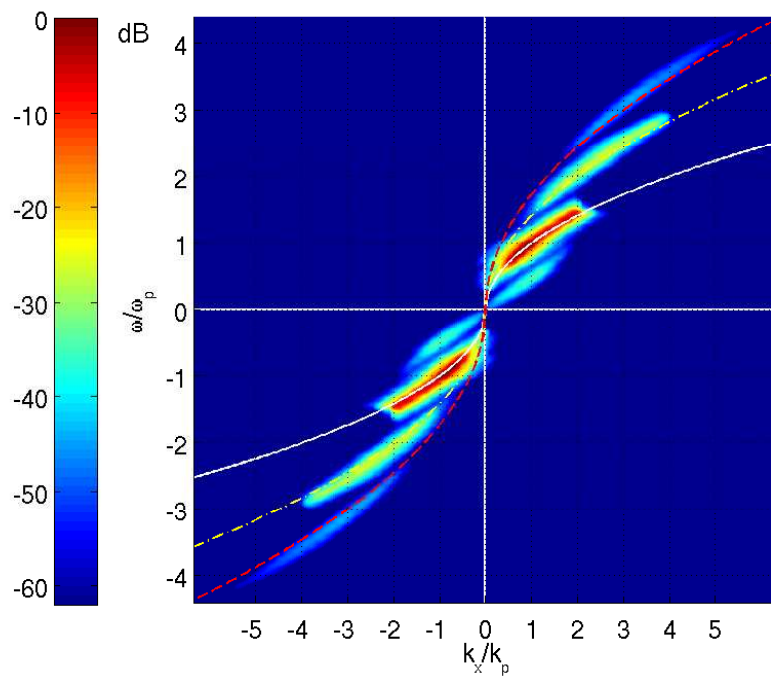


Figure 2: MNLs: $k_x\omega$ -slice obtained from 3-D spectrum at $k_y = 0$. Color scale in dB relative to the peak. White solid curve, deep water dispersion relation; yellow dash-dot curve, second order contribution; red dashed curve, third order contribution.

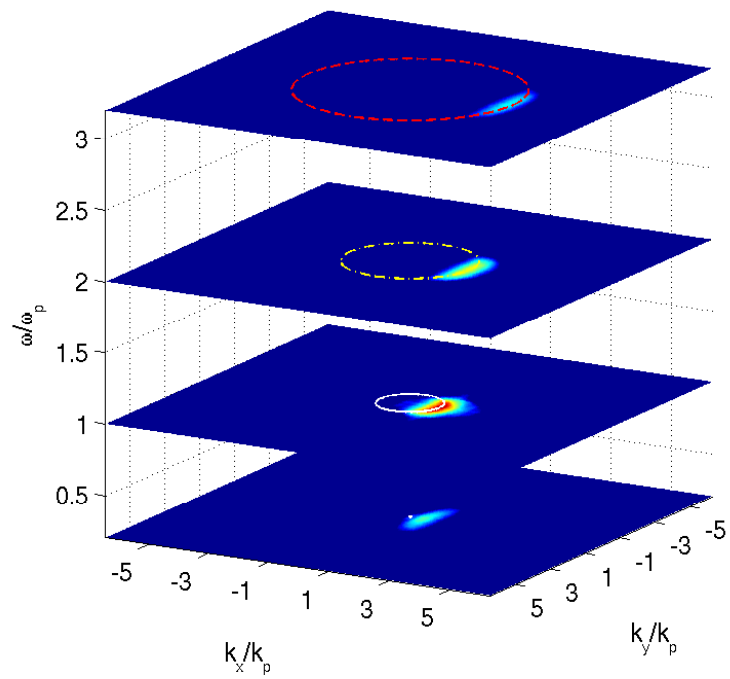


Figure 3: MNLS: Slices through (\mathbf{k}, ω) -spectrum showing $\phi_\omega(\mathbf{k})$ for constant frequencies. White solid circle, deep water dispersion relation; yellow dash-dot circle, second order contribution; red dashed circle, third order contribution.

200 data. In Fig. 4, $k_x k_y$ -slices of $\chi(\mathbf{k}, \omega)$, for selected ω , are shown. Since the resolution in the \mathbf{k} -plane is coarse, 40 neighboring slices have been averaged in the ω -direction. The result is heavily smeared and biased compared to the corresponding distributions obtained from simulated data (see Fig. 3). Nevertheless, the maximum distribution lies outside the dispersion circle for $\omega < \omega_p$ and within the dispersion circle for $\omega > \omega_p$.

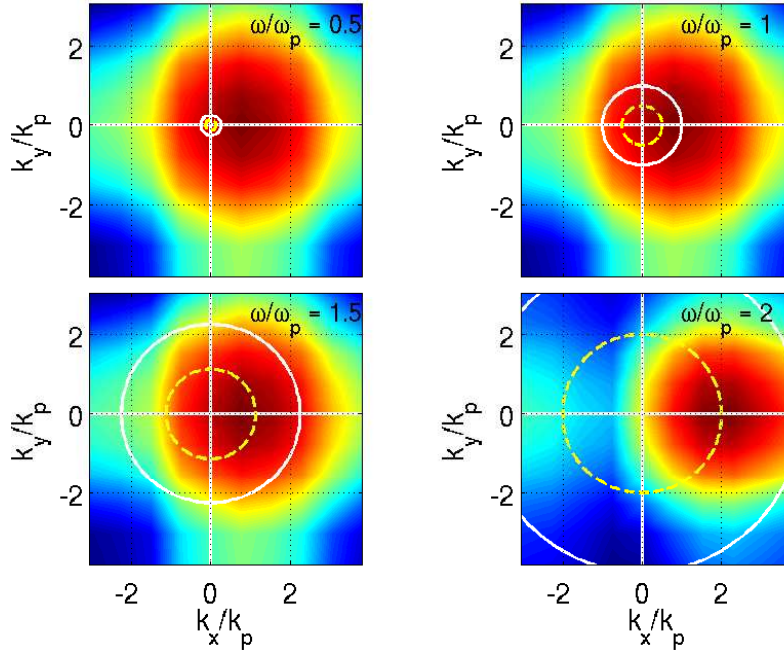


Figure 4: MARIN (location 2, full 3-D analysis, $\sigma_{1I} = 19.1^\circ$): $\phi_\omega(\mathbf{k})$ for fixed ω . All the plots are scaled with the maximum value of $\phi_\omega(\mathbf{k})$ for the given ω . Solid circle, linear dispersion relation at specific frequency; dashed circle, second order contribution.

205

From the estimate of $\chi(\mathbf{k}, \omega)$, one may obtain the moments for each frequency and then determine the directional parameters as discussed in Sec. 2.1. The data may also be analyzed using the standard ML and IML algorithms. These provide estimates of the directional distributions, $D(\theta, \omega)$, from which we then
 210 determine the directional parameters. Alternatively, the time series data may be

interpolated into HPR time series, and analyzed in that way, leading to estimates of the five basic moments. Figure 5 shows frequency spectra on the three sensors and estimated directional spectra, obtained by ML, IML and ME-B algorithms. The frequency spectra, in the upper left corner, are almost identical. Observe also the bend, slightly below twice the spectral peak, where the noise takes over the spectra. The spectra from the three algorithms are also very similar, with no distinct bi-modality which is perhaps masked by spatial aliasing in the high frequency region.

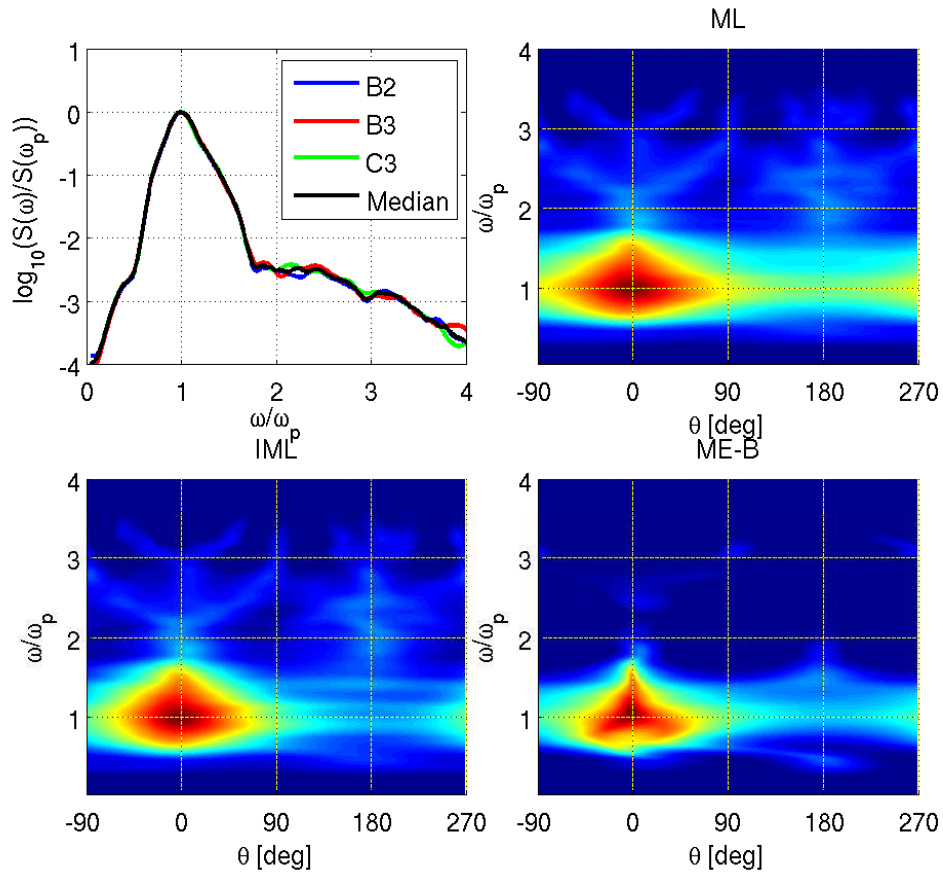


Figure 5: MARIN (location 2, array type 1, $\sigma_{II} = 19.1^\circ$): Frequency spectra on the three sensors and their median (upper left), ML, IML and ME-B directional spectra.

In Fig. 6, estimates of the directional spread from the three wave conditions
 220 are shown in the upper row. Even if the experiments are initialized by frequency
 independent directional distributions at the wave maker, the spectra develop
 directional spread strongly dependent on frequency with minimum spread near
 the spectral peak. For the narrow input spread ($\sigma_{1I} = 5.7^\circ$), the increase
 in the spread below the spectral peak is not so strong, and in this case, the
 225 increase in the check ratio toward the low frequency starts near the spectral
 peak. Increasing the input spread to $\sigma_{1I} = 12.5^\circ$ (medium) and $\sigma_{1I} = 19.1^\circ$
 (broad) changes the result moderately, in both cases, there is a strong increase in
 the spread below the spectral peak, however, there is no major difference in the
 check ratio. In all of the three cases, the directional spread increases smoothly
 230 from the spectral peak to about $\omega = 1.5\omega_p$ and then rapidly to maximum at
 about twice the spectral peak. The observed rapid increase to maximum is
 combined effect of spatial aliasing and lack of spectral coherence. The effect
 of spatial aliasing is also evident in the estimates of the mean directions in the
 lower row. The ML and IML directions veer off from the HPR directions above
 235 about two times the spectral peak. In these particular situations, there is a
 good agreement between the ML and Ekofisk spread around the spectral peak
 and for the three wave types while the agreement between Mitsuyasu and IML
 and HPR is best for the broader input spread.

Estimates of the directional spread by the standard ML-algorithm, around
 240 the peak spectrum, is too large compared to the HPR spread, whereas IML and
 HPR fit surprisingly well, in particular for the broader distributions.

Seven more datasets from each wave condition were analyzed to examine the
 spreading behavior more closely, in almost all of them, results were consistent
 and in agreement with the ones shown in Fig. 6.

245 When using array type 2, the ML results enhance significantly and become
 virtually similar to IML and HPR results on array type 1. The spatial devel-
 opment of the various wave parameters in the MARIN wave tank is shown in
 Fig. 7 for the broad spread ($\sigma_{1I} = 19.1^\circ$). In each plot, the graphs represent
 results from six locations along the propagation direction designated by 1 – 4,

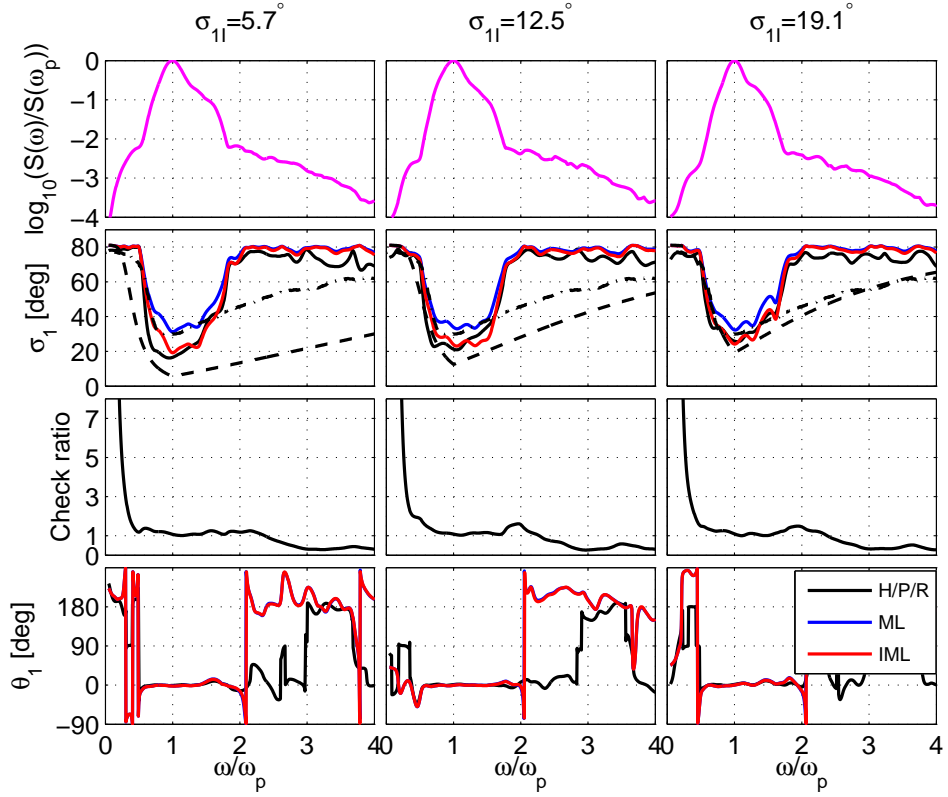


Figure 6: MARIN (location 2, array type 1): Frequency spectra (top row), directional spread (second row), check ratio (third row) and mean directions (bottom row) for narrow, medium and broad input spread. Black curve: standard triplet analysis (HPR); blue curve: ML analysis; red curve: IML analysis; dash-dot line: Ekofisk spread; dashed line: Mitsuyasu spread. The IML analysis is performed with $\omega_R = 0.2$ and 3 iterations.

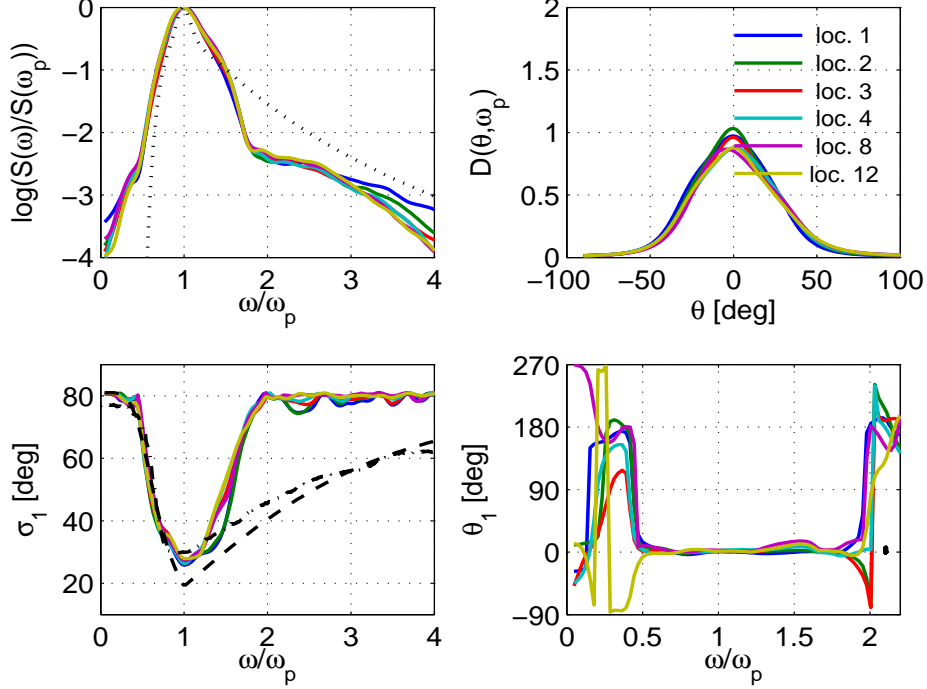


Figure 7: MARIN (ML on Array type 2, $\sigma_{1I} = 19.1^\circ$): Spatial evolution of frequency spectrum, directional distribution for peak frequency, directional spread, and the mean direction. Dotted line: input frequency spectrum; dash-dot line: Ekofisk spread; dashed line: Mitsuyasu spread.

250 8 and 12 in Fig. 1b, where the evolution distance between location 12 and 1 is about 14 peak wavelengths. Results are very similar, regardless of measurement locations, it is likely that the similarity observed is linked to a lack of spectral evolution. In the upper left corner, the input frequency spectrum is narrow compared to the estimate from the data. In the lower left corner, field direc-
 255 tional spread from the Ekofisk laser array is shown. It is an averaged spread from ten records with large H_s . The MARIN spread deviates from the Ekofisk spread at about 1.5 the spectral peak and increases to maximum at twice the spectral peak due to an effect of spatial aliasing and apparent lack of spectral coherence, see Sec. 3.1.

260 *4.2. Field observations*

We shall consider estimates of directional spread and distributions based on the ML, IML and HPR methods, all described in Sec. 2.2. The ML and IML algorithms provide estimates of the directional distributions from which we then determine $\theta_1(\omega)$ and $\sigma_1(\omega)$, whereas the HPR method provides the
265 four leading Fourier coefficients to, among others, the Burg-ME and Shannon-ME directional distributions. The Burg-ME produces two peaks when applied to the Fourier coefficients of, for instance, the $\cos-2s$ distribution. This peak splitting tendency is sometimes considered to be a weakness of the method [2]. We leave the details to a separate paper and concentrate here on illustrating
270 some of the results in Fig. 8 and 9. Frequency spectra of the three lasers (L) are almost identical and the agreement between the mean direction is perfect for the range of frequencies up to about four times the spectral peak. Above about four times the spectral peak, the ML and IML directions veer off from the HPR direction. This is a well known effect of spatial aliasing in the ML algorithms. Estimates of the directional spread are shown on the upper left corner, there
275 is a significant bias in the directional spread between ML and HPR estimates, whereas the IML and HPR spreads are surprisingly similar up to about four times the spectral peak. The HPR method gives the opportunity to study the estimator for the spread without making any assumptions about LWT. Besides,
280 it can be used as a benchmark to evaluate other directional analysis methods when possible.

For the directional spectra, the estimate by the standard ML differs significantly from both IML and ME estimates. The IML and ME spectra show more details compared to the rather smeared ML spectrum. Many ME ana-
285 lyzed data suggest a bimodal directional distribution for high frequencies, and in many situations, this bi-modality is even observed in the IML spectra for moderately high frequencies and in the Shannon based ME results for relatively higher frequencies. The highly different treatment for the data analysis of ME and IML suggests that in these situations, the bi-modality is real. However, in
290 many circumstances, the peaks in the Burg ME distributions are pronounced

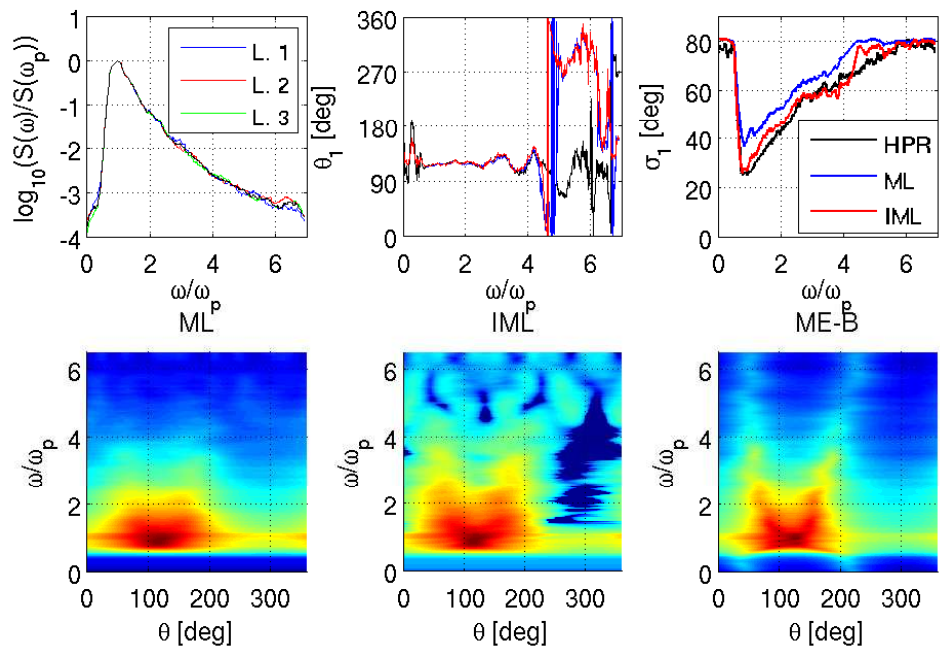


Figure 8: Ekofisk data recorded on the 12th of Oct. 2004 at 12:00, $H_s = 3.58\text{m}$. Upper left: Frequency spectra for the three input time series. Upper centre: Mean wave directions. Upper right: Directional spread. Lower row: Directional spectra.

compared to the moderate peaks observed in both IML and the Shannon based ME distributions (see Fig. 9).

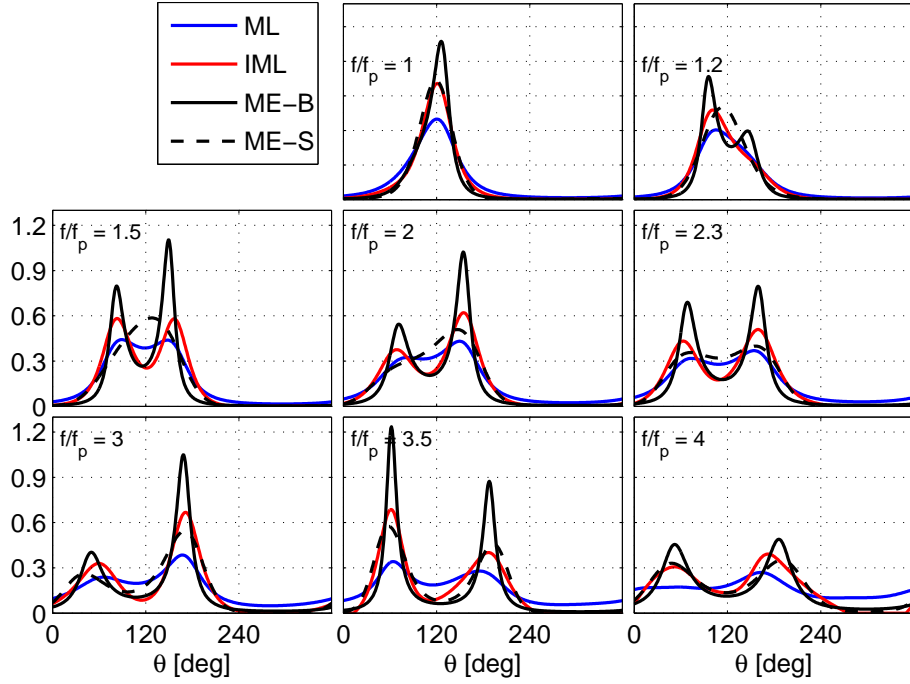


Figure 9: Ekofisk data recorded on the 12th of Oct. 2004 at 12:00, $H_s = 3.58\text{m}$. Directional distributions, $D(\theta, \omega)$, for selected frequencies.

4.3. Linear simulation with nonlinear reconstruction

In order to show that the reconstruction in Eq. 21, together with linear wave
 295 evolution, recovers the frequency dependent spread reported in [1], we present
 in Fig. 10 an HPR analysis of results from simulations carried out for infinitely
 deep water and with steepness $\epsilon = 0.05$ and spread $\sigma_{1I} = 5.7^\circ$, 12.5° and 19.1°
 similar to the MARIN experiments. In the simulations, only the linear part of
 Eq. 24 is integrated to discard the dynamic nonlinearity; on the other hand, the
 300 reconstruction in Eq. 21 takes the static nonlinear contributions into account.
 The estimated directional spread is different in shape from the input to the

model, it is strongly dependent on frequency and remains the same everywhere in the numerical basin.

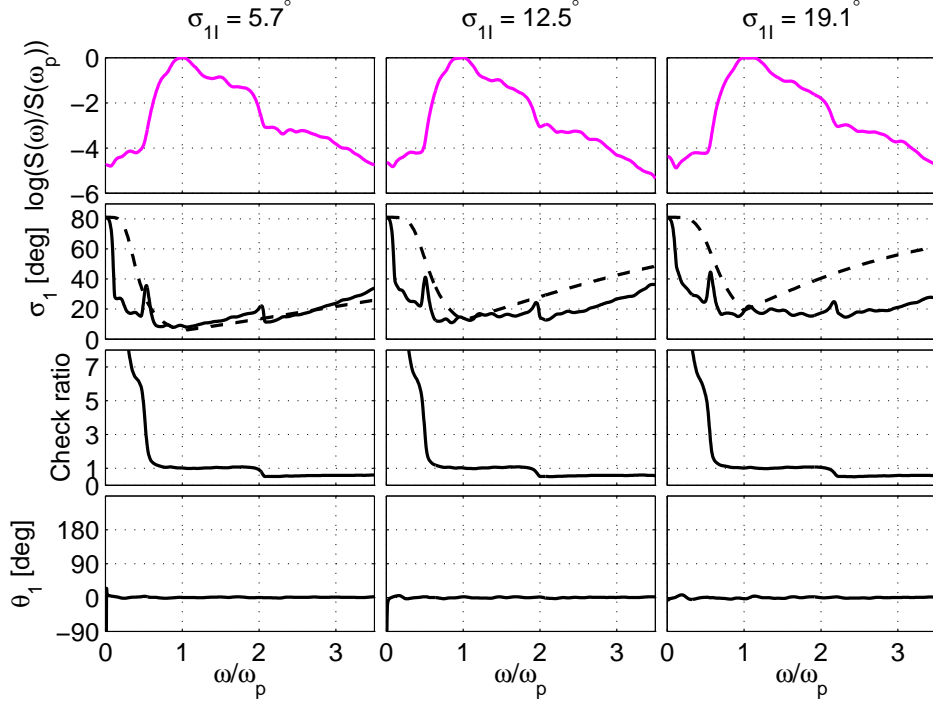


Figure 10: Simulation with the linear part of Eq. 24 together with reconstruction according to Eq. 21: Similar to Fig. 6 with the HPR method alone. Dashed line: Mitsuyasu spread.

For the narrow distribution ($\sigma_{1I} = 5.7^\circ$), the spread clearly increases towards
 305 both high and low frequencies, whereas, for broad distribution ($\sigma_{1I} = 19.1^\circ$) the
 increase toward the high frequencies is quite weak. In all of the above cases, the
 minimum spread lies below the spectral peak. The check ratio, in the middle
 row, shows a strong increase towards the low frequencies and a drop from 1 for
 frequencies higher than twice the spectral peak, where the second order spectral
 310 contribution is expected to dominate.

The numerical experiment has been checked by not including $\bar{\eta}$ in the recon-
 struction equation, Eq. 21. The resulting spread and check ratio did not show
 frequency dependent variations demonstrating that it is primarily the contribu-

tion from $\bar{\eta}$ that produces the directional spread reported in [1]. Except for very
 315 wide spectrum the agreement between our results and results in [1] is quite good.
 The observed behavior in the spread and check ratio, below the spectral peak,
 is therefore originated from second order difference-frequency contributions.

4.4. Nonlinear simulation

Finally we show numerical simulations using the full eqs. (21)–(24) demon-
 320 strating that the establishment of frequency dependent directional spread is not
 only a matter of static nonlinear reconstruction. In Fig. 11, the input direc-
 tional spread has been set to 5.7° , 12.5° and 19.1° , and the steepness has been
 set to 0.05 and 0.1. With the dynamic nonlinearity in the evolution equation,
 the directional spread is found to grow with an increase in input wave steepness,
 325 the minimum spread is also moved toward the spectral peak. This is shown even
 more clearly in Fig. 12 which employs six different steepnesses between 0.05 to
 0.1, inclusive.

Depending on the steepness of the waves, the numerical simulations demon-
 strate the necessary propagation distance required for sufficient development
 330 of the spread. As seen in Fig. 11, for $\epsilon = 0.05$ the directional spread, $\sigma_1(\omega)$,
 does not change significantly with evolution distance. This result is supported
 by the observations from the MARIN experiment. For $\epsilon = 0.1$, the directional
 spread develops over a distance of 80 peak wavelengths before it becomes nearly
 stationary. We thus draw the conclusion that the distance for the frequency de-
 335 pendent directional spread to become fully established increases as the steepness
 increases.

From Fig. 11 it can be appreciated that the crest length decreases with
 increasing propagation distance. From Fig. 12 it can furthermore be appreciated
 that this effect is enhanced with increasing input steepness.

340 5. Conclusion

We have found both experimental and numerical evidence that a frequency
 dependent directional spread naturally develops from an initial state of no fre-

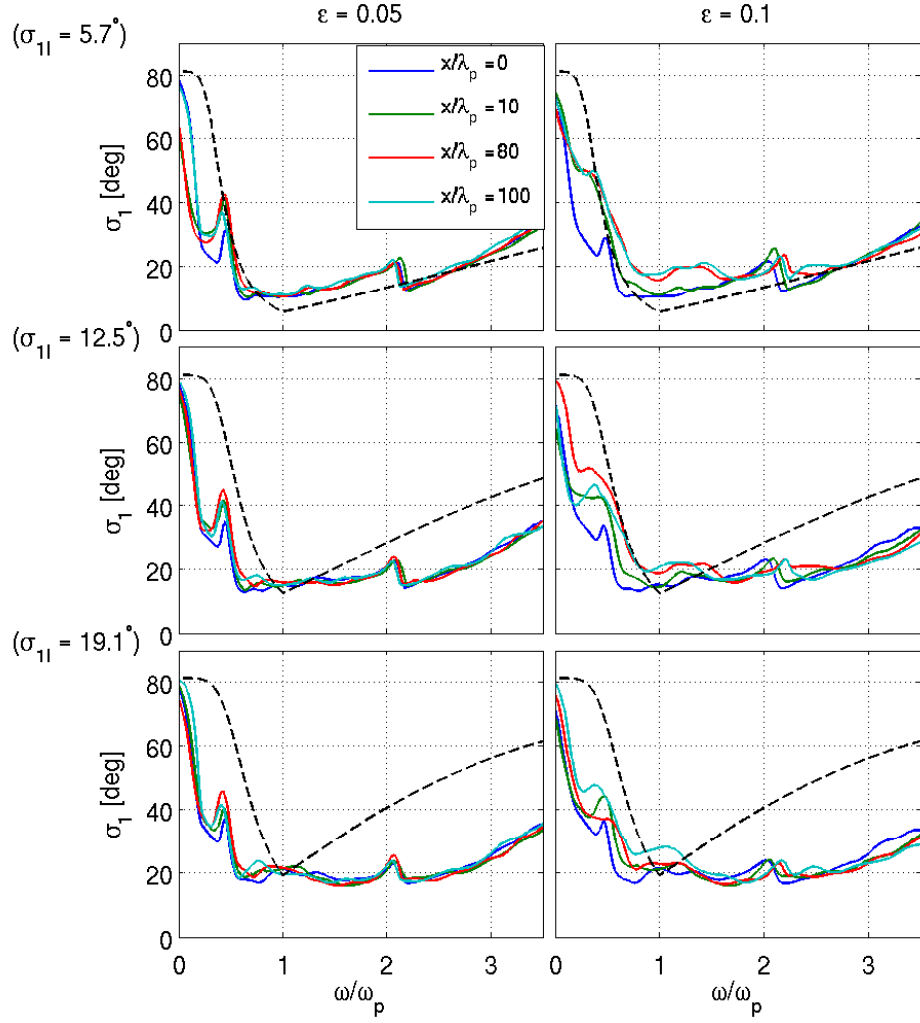


Figure 11: MNLS: Effects of dynamic nonlinearity and wave steepness on the development of the directional spread for $\sigma_{1I} = 5.7^\circ$, 12.5° and 19.1° ; dashed line, Mitsuyasu spread.

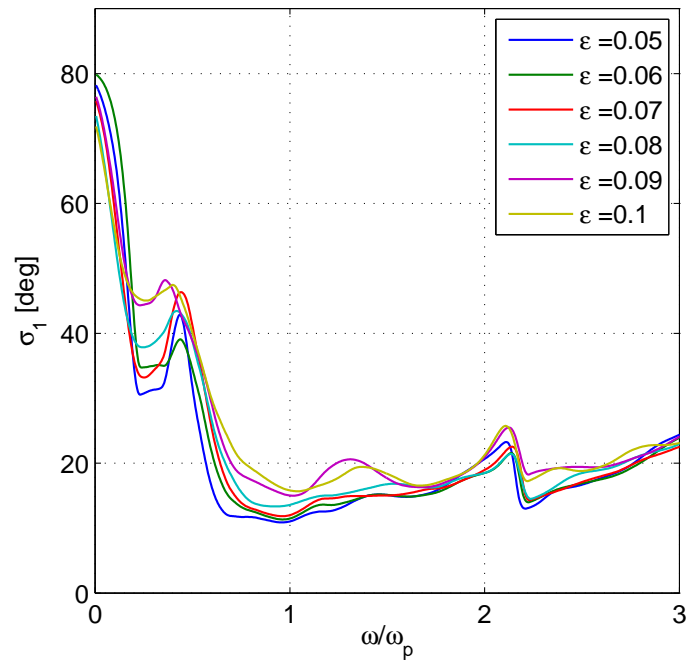


Figure 12: MNLs: Directional spread at 100 peak wavelengths from the numerical wave paddle. Input spread $\sigma_{1I} = 5.7^\circ$ and steepness $\epsilon = 0.05, 0.06, 0.07, 0.08, 0.09$ and 0.1 .

quency dependence. The frequency dependence is partly a consequence of static nonlinear contributions which act instantaneously, and partly a consequence of
345 dynamic nonlinear contributions which need some evolution distance to develop. The static contribution occurs primarily below spectral peak, the dynamic contribution occurs both above and below spectral peak. The necessary distance for developing a steady frequency dependent directional spread seems to increase with increasing steepness of the wave field.

350 An important consequence of these observations is that it is not necessary to generate a frequency dependent directional spread in wave basins since the wave field will tend to develop an adequate frequency dependent directional spread by itself, care just needs to be taken that the wave field has been allowed to develop over a sufficiently long distance that a steady frequency dependence has
355 been established.

The analysis of datasets from the Ekofisk field indicates that a large fraction of the directional wave spectra have a bimodal character above the spectral peak. The bi-modality is observed for all four algorithms employed in the paper suggesting that the bi-modality is not limited to, as often claimed, an artifact
360 of the Burg ME algorithm.

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365 Kongsberg, SBM Seaflex, Sirehna, Statoil and Total) for making the data available for this study. The field data are from the laser array at the Ekofisk field in the North Sea and used by permission of ConocoPhillips Inc.

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